

# Spatio-temporal evolution of temperature and fluid flow through a new “thermo-lithological” boundary; the case of a pit crater of Karthala volcano (Comoros archipelago) refilled on January 13th 2007 by a lava flow

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## ABSTRACT

On January 13th 2007 an effusive eruption on the top of Karthala volcano, in the Comoros islands, emitted a lava flow which has been perfectly constrained inside the cylindrical shape of the *Choungou Chagnoumeni* pit crater; 225 m in diameter and with a thickness of about 7 m.

This eruptive event, with a known geometry before the eruptive event, give a unique opportunity to study a “thermo-lithological” structural boundary recently born, and follow its thermal and fluid flow evolution in time. The motivation of this survey is to better understand some scientific questions such as: How fluid flows evolve in time along a new thermo-lithological boundary? How lava temperature cooling induce changes in hydrothermal circulation? Different data set have been acquired on the field in order to discuss these problematics. On March 2008, a temperature monitoring was installed with a data logger (CR 1000 Campbell Scientific) and 11 sensors placed at 30 cm depth, 1 m apart, inside the crater, along a straight profile, perpendicular to the western rim of the *Choungou Chagnoumeni* crater. Data were recorded during 275 days out of 677 days, through 4 different periods between 3rd March 2008 and 8th January 2010.

On June 2009, high resolution electrical resistivity tomography (ERT) coupled with soil CO<sub>2</sub> diffuse degassing was performed above the monitoring line of temperature sensors along a profile of 128 m long, with ERT electrodes and CO<sub>2</sub> measurements located each 2 m.

Numerical modeling of the temperatures was also tested on this case study, thanks to a finite element method (FEM) code using the *RheolefC++* library. The aim of using a numerical code, integrating a maximum of parameters, have been to better interpret the other data acquired on the field.

The lateral permeability transition induced by the crater refilled with lava flow material was attested by the soil diffuse degassing profile.

Coupling ERT and temperature modeling allow evidencing the heat transfer with conductive processes.

The most striking result has been the inversion of the lateral temperature gradient which evolves during the first three years of cooling, from the highest temperature in direction to the center of the crater to the highest temperature in direction to the boundary of the crater.

Such a result highlights the importance of structural boundary in dragging hydrothermal fluids and has been interpreted by a change in the main parameter governing the heat transfer close to the boundary of the crater, which evolved from (1) a lava heat source toward (2) a higher permeability area constituted by the boundary of the crater. This study clearly displays the importance of the lateral permeability gradient in changing the hydrothermal circulation in time.

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## 1. Introduction

### 1.1. General context

On active volcanoes, temperature appears as one of the most intuitive parameter to look for precursors of major or paroxysmal eruptive

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activity. This parameter can be measured directly by temperature sensors inserted into the soil (Friedel et al., 2004; Ricci et al., 2015; Gaudin et al., 2017) or by remote sensors (Wright et al., 2004; Coppola et al., 2015a; Laiolo et al., 2017). Thermal anomalies recorded on active volcanoes can be subdivided into two groups depending on how the heat transfer is directed toward the surface:

First, the anomalies related to hydrothermal fluid flow condense in the atmosphere or close to the surface (Aubert, 1999; Aubert et al., 2008). These are of low amplitude, generally reaching just the threshold of the water ebullition point. These hot fluids are generally underlining structural discontinuities with higher permeability such as crater or caldera rims or tectonic faults, along which fluids flow preferentially (Finizola et al., 2003, 2006; Barde-Cabusson et al., 2009a; Antoine et al., 2009; Revil et al., 2004, 2008, 2011; Peltier et al., 2012). Sometimes, temperatures are much higher (several hundred of degrees), because gases escaped from magma at depth and reach the surface without temperatures be buffered by hydrothermal systems (Chiodini et al., 1995).

Second, anomalies of higher amplitude, generally above 100 °C, are directly related to lava flows or lava lakes (Wright et al., 2001; Harris et al., 1999, 2002, 2005; Coppola and Cigolini, 2013; Coppola et al., 2015b).

While hydrothermal origin anomalies expand generally over few tens up to few hundred meters large (Finizola et al., 2002; Bennati et al., 2011), magmatic origin anomalies can extend over larger distance (Harris et al., 1997; Barde-Cabusson et al., 2009b; Coppola et al., 2009). Sometimes, interaction between these two systems occurs; because structural boundaries are also higher permeability areas where magma can migrate toward the surface.

In our study case, on Karthala volcano, in Comoros islands, the eruption of 13th January 2007 and its lava flow that refilled a near perfect cylindrical volume constituting the *Choungou Chagnoumeni* crater, give a unique opportunity to study a hot structural boundary recently born, and follow its thermal and fluid flow evolution in time.

The motivation of this survey is to better understand scientific questions such as:

How fluid flow evolve in time along a new “thermo-lithological” boundary?

How lava temperature cooling induce changes in hydrothermal circulation?

Such problematic are interesting not only in volcanology in term of time evolution in fluid circulation and self-sealing along a structural boundary (Finizola et al., 2009; Revil et al., n.d.), but also in mining geology for related mineralization induced by magmatic/hydrothermal fluid flow (Hedenquist and Henley, 1985; Hedenquist and Lowenstern, 1994).

In this study, soil heterogeneity and spatial distribution of relative lateral changes in soil permeability is assessed coupling soil gas technique and electrical resistivity tomography (Finizola et al., 2006; Revil et al., 2011).

Numerical temperature modeling is used to locate the spatio-temporal evolution of temperature imposed by the lava cooling. Considering that the distribution in space of the true permeability values is not known, we decided to not introduce in the model parameters strongly affected by permeability variations; as a consequence, convection at depth have not been considered in the model. Details of the different parameters taken into account in the numerical temperature modeling are specified in the following chapters.

Temperature monitoring at 30 cm depth have been installed along a straight line of 10 m long, with 1 m apart, perpendicularly and inside the crater rim, to study (1) the general temperature cooling in time, (2) the lateral influence of the structural boundary, and (3) the location of the most active convective area and its evolution in time.

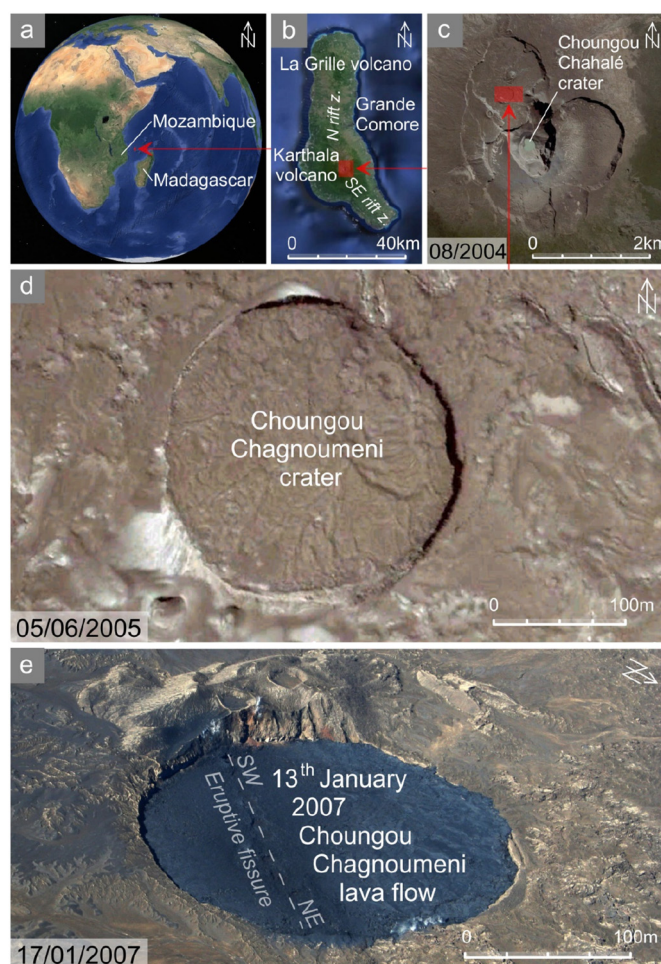
## 1.2. Geological framework

Karthala (2361 m a.s.l.) is an active basaltic shield volcano located on *Grande Comore* Island, the youngest and the most westerly of the four

volcanic islands of the Comoros Archipelago (Bachèlery and Hémond, 2016). Located between the northern Madagascar and Mozambique (Fig. 1a), this alkali volcanism constitute a possible extension of the East African rift (Michon, 2016). Petrology of Karthala volcano is mainly constituted of basalt and few hawaiiite (Bachèlery and Hémond, 2016).

The Island of *Grande Comore* is subdivided into two main quaternary volcanic edifices; (1) La Grille volcano, the oldest one located on the north, and (2) the Karthala volcano, forming the southern two-third of the island (Fig. 1b). The shape of the island is underlining the two main rift zones (North and South East) crossing the volcanic edifice. The summit part of Karthala is characterized by a polylobate caldera complex approximately elliptical (3.5 × 2.8 km) (Strong and Jacquot, 1970; see Fig. 1c). The eruptive style is mostly effusive along the rift zones, with eruptions along eruptive fissures, but also far from the rift zones at low elevation, or inside the summit caldera (Krafft, 1982; Bachèlery and Coudray, 1993). However, the non-permanent summit water lake located inside the main crater (*Choungou-Chahalé* crater; see Fig. 1c) is also responsible of phreatic or phreatomagmatic eruptions, whose ash deposits affect generally all the *Grande Comore* Island (Bachèlery et al., 1995; Bachèlery et al., 2016; Morin et al., 2016).

Since the beginning of the 19th century, a total 39 eruptions have been recorded but of which 14 remains uncertain. Karthala volcanic activity can be assimilated to a typical shield volcano activity, constituted



**Fig. 1.** (a) Location on Google Earth of Comoros archipelago, between Mozambique and Madagascar; (b) Google Earth picture of *Grande Comore* island; (c) Google Earth picture of the summit calderas of Karthala volcano; (d) Google Earth picture of the *Choungou Chagnoumeni* crater before the January 13th 2007 eruption (e) Aerial picture of the *Choungou Chagnoumeni* crater after the January 13th 2007 eruption (courtesy of P. Bachèlery).

mainly of eruptive fissures, lava flows of pahoehoe and a type and scoria cones, sometimes disrupted by phreatic or phreato-magmatic events.

This recent eruptive activity of Karthala can be subdivided into three main periods where the eruptive activity was located geographically in three different areas:

- 1) During the first one, between 1848 and 1903, the activity occurred both on the south east and north rift zones, with eruptive fissures located on the axis of the rift zones associated with lava flows between 8 and 13 km long, reaching the eastern coast (1848, 1858, 1862, 1876 and 1880) or the western coast (1858 and 1859). Only one lava emission occurred outside of the rift zone (1872), at the north-western foot of Karthala volcano, at about 250 m a.s.l., and also reach the western coast close to Vanamboini village.
- 2) During the second period, between 1904 and 1972, lava emission was localized only along the north rift zone (1904 and 1918), but the associated lava flows do not reached the sea as in the previous period. Eruptive activity also occurred inside the summit caldera, with a phreatic eruption inside the main eruptive crater, the *Choungou Chahale* (1918). The other eruptions were purely magmatic and located inside the main crater or inside the northern lobe of the caldera, constituting the extension of the north rift zone (see below). The 1977 eccentric magmatic eruption, near the western coast, with eruptive fissures at about 360 m a.s.l., and associated lava flows destroying part of Singani village, constitutes a transition toward the third period.
- 3) During the third period, between 1991 and the present, the activity was localized only in the summit caldera. A succession of three phreatic eruption, affecting in both cases a large part of Grande Comores inhabitants (evacuation of people by local authorities, water tank contamination, health problems due to the fine ash-falls, panic among the villagers; see Morin et al., 2016, Chapter 25), occurred in the *Choungou Chahale* crater, in July 1991, April 2005 and November 2005.
  - (a) The eruption of 11th July 1991 was purely phreatic and proceeded by 7 days of intense seismicity (4000 earthquakes per day) and ended in <24 h.
  - (b) Before the phreato-magmatic eruption of 16th April 2005, an increase in the seismicity also occurred during the month of March with 40 earthquakes on average per day. The eruption of April 2005 lasted with an activity of active lava lake started on 17th April and ended on 19th April.
  - (c) At the end of August 2005, another period of greater seismicity began and resulted in another phreato-magmatic eruption on 24th November 2005. This eruption evolved quickly (25th November) in a lava lake activity which ended on 8th December 2005 with two small cones with strombolian activity. This succession of three explosive eruptions, with a phreatic component, was triggered in both cases by the water lake present in the bottom of *Choungou Chahale* crater.

Thanks to the November 2005 eruption, the effusive activity inside *Choungou Chahale* crater was enough in term of thickness of lava lake to avoid the reappearance of the water lake at the bottom of the crater. As a consequence, the following eruption, which occurred inside the *Choungou Chahale* crater on 28th May 2006, was purely magmatic. This eruption began after one week of higher seismicity (20 earthquakes per day between 20th–27th May 2006), was characterized by huge lava fountaining activity and ended on 31st May 2006.

The last eruptive activity at Karthala volcano occurred on 13th January 2007, after an earthquake ( $M > 4$ ) on 12th January 2007, felt by Grande Comores inhabitants. This eruption that lasted <24 h, was purely magmatic and located only inside the *Choungou Chagnoumeni* crater (Bachèlery et al., 2016).

During historical times, the most active area of the summit caldera was inside the *Choungou Chahale* crater (1918, 1948, 1952, 1959?, 1991,

April 2005, November 2005, May 2006; see Fig. 1c). Outside this crater, the other area of high level of activity is the northern lobe of the caldera associated with the most intense hydrothermal circulation (Lénat et al., 1998) and seismicity (Savin et al., 2005). Moreover, some eruptions occurred in this area; lava flow in 1965, scoria cones and aa lava flow in 1972 covering nearly the entire northern lobe of the caldera and overflowing to the north, and four eruptions of the *Choungou Chagnoumeni* pit crater (1918, 1948, 1965, and 2007), which is qualified for that of second active crater of Karthala. This latter, of about 225 m in diameter, with vertical walls, was formed during the 1918 phreatic eruption, which structured both the northern part of the present-day *Choungou Chahale* crater and the *Choungou Chagnoumeni* crater. From (1) a depth of about 150 m in 1936 (after its formation in 1918), the *Choungou Chagnoumeni* crater was gradually infilled to (2) a depth of about 50–60 m in 1952 during a magmatic activity of *Choungou Chahale* affecting *Choungou Chagnoumeni* pit crater, and (3) about 30 m depth after the 1965 *Choungou Chagnoumeni* lava flow activity, and (4) only 7–8 m depth (7,26 m exactly with average calculation from DEM's; Nicolas Villeneuve, personal communication) after the large magmatic eruption of 1972 refilling large part the northern lobe of the summit caldera (Bachèlery and Coudray, 1993; see Fig. 1d). The crater was the site of the 13th January 2007 eruption, and was completely infilled in <24 h by a volume of about 275,000 m<sup>3</sup> of basaltic lava emitted from a NE-SW eruptive fissure located inside the *Choungou Chagnoumeni* crater (Bachèlery et al., 2016). A line of spatter cones marks the eruptive fissure crossing the *Choungou Chagnoumeni* crater (see Fig. 1e). This lava can be likened to a rough surface pahoehoe lava flow. In this paper, we will use “t<sub>0</sub>” to talk about of the day of the eruption (13th January 2007). During the first mission in March 2008, the thickness of post-eruption sedimentary deposits, brought mainly by rainwater, in the studied area, was of about 25–30 cm, constituted of fine ash removed material (Fig. 2a, c). This fine sedimentary deposit allows performing good quality temperature and soil gas measurements, limiting both the cooling and the dilution with the air. Without this thickness of sedimentary deposits, temperature and soil gas measurements above the 13th January 2007 rough pahoehoe lava flow would have been impossible, due to the too strong interaction (temperature and chemical dilution) with the air.

## 2. Methods

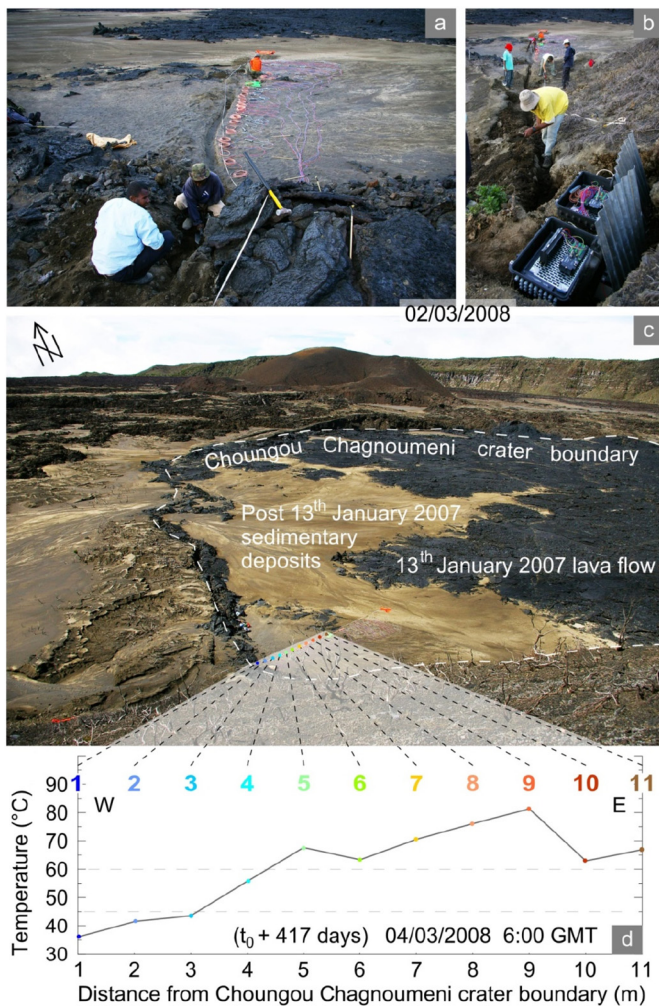
### 2.1. Data acquisition and processing

At t<sub>0</sub> + 420 days (in March 2008), a Campbell Scientific datalogger (CR1000) with a multiplexor (AM2 5 T) was installed close to the *Choungou Chagnoumeni* crater. After digging a trench, 11 temperature sensors (thermocouples K type) were placed at 30 cm depth along a straight profile, perpendicular to the western rim of the *Choungou Chagnoumeni* crater, beginning inside and at 1 m from the border toward the center of the crater (Fig. 2a, c). The spacing between each sensor is 1 m. The CR1000 recorded 1 measurement per hour for each sensor. As a safety reason for the scientific equipment (vandalism), no solar panel has been installed and all the equipment was hidden under the ground (Fig. 2b). As a consequence, the data set was obtained during 4 periods of time, when the battery was changed:

- (1) From 3rd March 2008 to 5th May 2008 (during 64 consecutive days),
- (2) From 22nd August 2008 to 23rd October 2008 (during 63 consecutive days),
- (3) From 15th June 2009 to 18th August 2009 (during 65 consecutive days),
- (4) From 18th October 2009 to 8th January 2010 (during 83 consecutive days).

Hence, during the period 3rd March 2008–8th January 2010, 275 days out of 677, corresponding to >40% of this period, was





**Fig. 2.** (a) Installation of the eleven temperature sensors at 30 cm depth inside the Choungou Chagnoumeni crater; (b) Installation of the Campbell Scientific monitoring station outside the Choungou Chagnoumeni crater; (c) Large view picture showing the location of the eleven temperature sensors located inside the Choungou Chagnoumeni crater; (d) Temperature of the eleven sensors measured at 30 cm depth, two days after the installation. The dashed lines link the bottom temperature plot abscissa (figure d) to the position of the sensors on the photography (figure c).

monitored. No data was lost during the consecutive days of monitoring period when the battery was working.

On 4th March 2008 at 6:00 GMT, the second day of monitoring (corresponding to  $t_0 + 417$  days), we can note at 30 cm depth a clear lateral temperature gradient from  $<40$  °C up to  $>80$  °C, increasing globally toward the center of the crater (Fig. 2d). This result is in good accordance with the original source of heat; the lava flow located inside the Choungou Chagnoumeni crater.

Moreover, at  $t_0 + 884$  days (in June 2009 and during the third monitoring period), we performed above the monitoring sensors, an electrical resistivity tomography (ERT) survey coupled with soil CO<sub>2</sub> concentration measurements.

ERT data were obtained using a set of 64 stainless steel electrodes with 2 m spacing between the electrodes, the ABEM Terrameter SAS-4000 resistivity meter and ES1064 multiplexer. The contact of the electrodes with the ground was improved by adding salty water.

A unique profile of 126 m long (63 spacings between 64 electrodes with 2 m spacing between take-outs) was performed positioning the middle of the profile, just on the western rim of the Choungou Chagnoumeni crater (Fig. 2c).

A Wenner alpha array was setup both for the good signal-to-noise ratio and the good sensitivity for the detection of vertical changes in

the subsurface resistivity below the centre of the electrode array. For each acquired measurement, we performed stacking to get a standard deviation better than 5% with a maximum of 4 stacks.

A good contact resistance between electrodes and medium allowed injecting a maximum of 200 mA current.

We acquired a total of 472 measurements of apparent resistivity. Topography is totally flat along our profile. The ERT data were then inverted by means of the commercial package RES2DINV (Loke and Barker, 1996) using a finite element grid for the forward modeling of the apparent resistivity. The RMS error obtained after 5 iterations is 6.1%. Resistivity tomogram is shown in Fig. 3b.

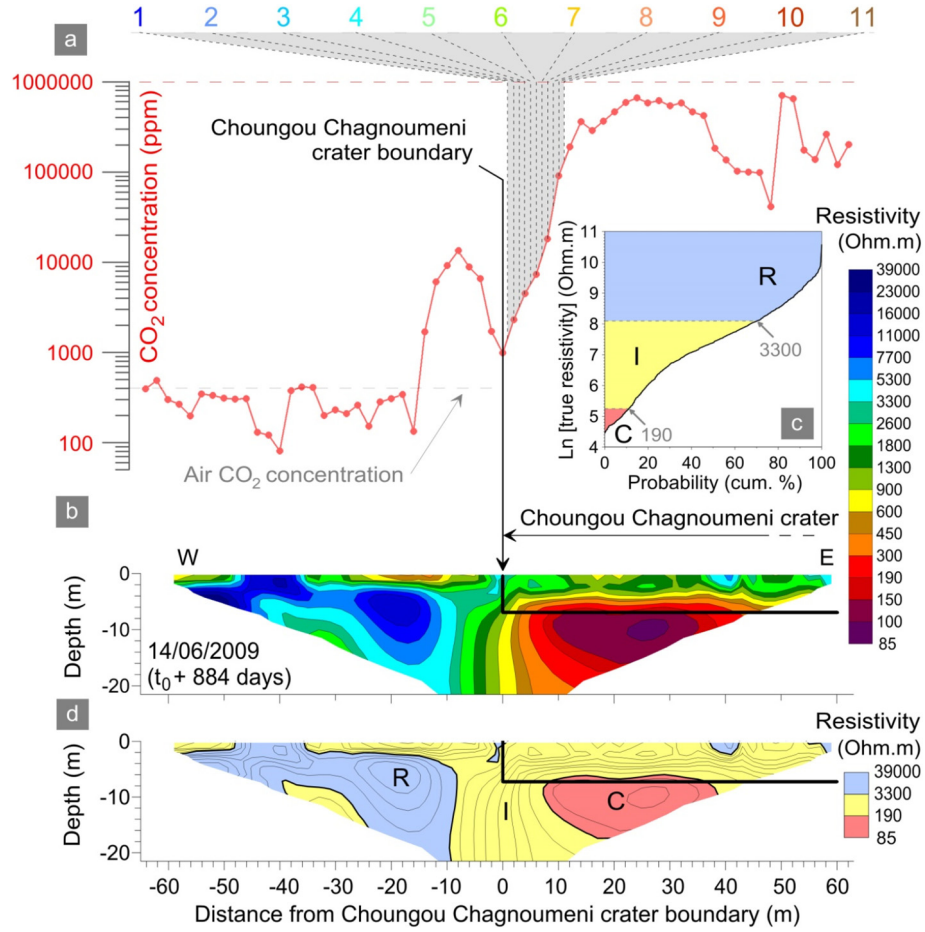
In addition to 2-D DC-resistivity tomography, we also acquired contemporaneously, soil CO<sub>2</sub> concentration measurements. These measurements were obtained with a spacing of 2 m along the profile, taking measurements in correspondence with all the ERT electrodes.

Soil gas was sampled through a copper tube (2 mm in diameter). This copper tube was first inserted in the soil to a depth of 0.5 m. The gas was analyzed directly in the field by infrared spectrometry (Edinburgh Instruments, model GasCheck). The analytical uncertainty was 5% of the concentration value (see Revil et al., 2011). The CO<sub>2</sub> concentration profile is shown in Fig. 3a.

## 2.2. Numerical method

The cooling of lava flow has been modeled in different context. Shaw et al. (1977) produced one of the first numerical formulations to model the cooling of lava lakes. They solved the heat balance equation in one-dimension by a method of explicit cell balances using a stack of elements. A temperature of 0 °C is fixed at the top of the stack and 1000 °C is given for the initial temperature of magma. The heat loss occurs by conduction and by adding two source functions for the latent heat of crystallization and for the vaporization of the rain. Peck et al. (1977) applied these numerical formulations on the specific case of the Alae lava lake from Hawaii which provides good results compare to the observations. Thereafter, many authors have studied and quantified by analysis and observations the main heat transfer mechanisms with the atmosphere: radiation, natural convection and forced convection by the wind (see Head and Wilson, 1986; Keszthelyi, 1995; Keszthelyi et al., 2003). Keszthelyi and Denlinger (1996) have developed a numerical model for the initial cooling of pahoehoe lobes taking into account the conduction, the thermal radiation, the forced and free convection and the latent heat of crystallization. Their formulation is based on a fully explicit finite-difference algorithm. Neri (1998) has developed another formulation based on an original application of the integral heat balance method (Goodman, 1964; Ozisik, 1993). The model involves radiation, forced or natural convection. It also taking into account the crystallization and the formation of crust layer at the surface where the thermal properties can be different than the mushy layers one. Equations are numerically solving by ordinary differential equations, in terms of surface temperature and by a Runge-Kutta method in the crust and mushy layer thicknesses. Other articles deal with the effect of the cooling on the dynamic of a lava flow (see Danes, 1972; Dragoni, 1989; Harris and Rowland, 2001; Bernabeu et al., 2016; Chevrel et al., 2018).

Note that all codes presented in Section 2.2 which all based on an explicit time schemes and uniform spatial grids. From one hand, we choose here to use an implicit scheme, since it is more robust for long time simulations. From other hand, we choose to base our computations on non-uniform spatial grids, suitable to take into account a large substratum region. So, we decided to write a code based on the finite element C++ library Rheolef for our numerical model of the Choungou Chagnoumeni cooling. Using the cylindrical symmetry of the geometry, the calculation domain  $\Omega$  is reduced to a vertical radial plane cut (see Fig. 4). It is composed in two parts: the crater occupied by the lava with a radius  $r$  and a depth  $H$  and the substratum. To avoid the boundary



**Fig. 3.** (a) Soil CO<sub>2</sub> concentration profile crossing the rim of the *Choungou Chagnoumeni* crater; (b) Electrical resistivity tomography (ERT) profile crossing the rim of the *Choungou Chagnoumeni* crater; (c) Probability plot analysis (Sinclair, 1974) of the true resistivity values measured on the ERT profile; (d) Electrical resistivity tomography (ERT) profile showing the three ranges of resistivity values identified through the probability plot technique. “R”, “I” and “C” stand for Resistant, Intermediate and Conductive body, respectively.

effect, we choose for the numerical simulation a substratum region with a size of 20 times the thickness of the deposit (145.2 m). The evolution of the temperature  $T$  inside  $\Omega$  is model for all time  $t > 0$  by the heat

equation using the Fourier's law conduction and a source term coming from the latent heat of crystallization:

$$\rho C_p \frac{\partial T}{\partial t} - \text{div}(\lambda \nabla T) = L \frac{\partial \phi}{\partial t} \text{ in } \Omega,$$

where  $\rho$  is the lava density considering the porosity and determined from the densities of the various lava flow levels (see Deroussi et al., 2009),  $C_p$  the heat capacity,  $\lambda$  the diffusion,  $L$  the latent heat of crystallization and  $\phi$  the crystallinity (i.e. the solid volume fraction). The numerical value of  $L$  has been taken in Chevrel et al. (2018).

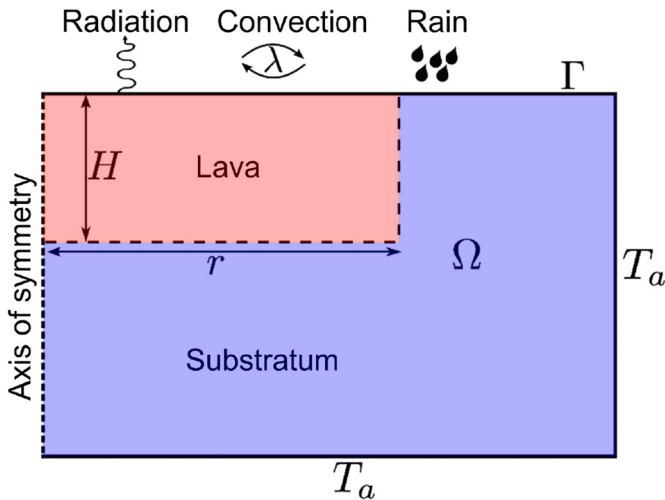
We assume as in Huber et al., 2009 that the crystallinity depends only of the temperature by the relationship

$$\phi = \frac{T_{liq} - T}{T_{liq} - T_{sol}} \text{ for } T_{sol} < T < T_{liq},$$

$$\phi = 1 \text{ for } T < T_{sol} \text{ and } \phi = 0 \text{ for } T > T_{liq},$$

where  $T_{liq}$  is the liquidus temperature and  $T_{sol}$  the solidus temperature. The value of  $T_{liq}$  has been taken in Dragoni, 1989 and the value of  $T_{sol}$  in Chevrel et al., 2018.

The model incorporates at the interface  $\Gamma$  between  $\Omega$  and the atmosphere the radiative flux, the convective flux and the heat flux to warm up and vaporize the rain (see Fig. 4). Indeed, we suppose that the vaporization of rain is only located in surface and we neglect the water infiltration, the vaporization in depth and the rise of the vapor. Thereby, when the surface temperature is less than the vaporization temperature



**Fig. 4.** Representation of the geometrical model used for the numerical modeling with the FEM code written with *Rheolef*: the calculation domain  $\Omega$  occupied by the lava and the substratum at the center; the interface  $\Gamma$  between  $\Omega$  and the atmosphere with the different heat flux on the top; the axis of symmetry on the left and the ambient-imposed temperature condition at the bottom and on the right.

(100 °C), the rain has no more effect on the cooling except the water heater. The air conduction (0,0262 W.m.K; see Adrian and Allan, 2003) is extremely small and will be neglected.

$$-\kappa \frac{\partial T}{\partial n} = q_{rad} + q_{conv} + q_{vap} \text{ on } \Gamma,$$

where:

$$q_{rad} = \varepsilon \sigma (T^4 - T_a^4),$$

$$q_{conv} = h_c (T - T_a),$$

$$q_{vap} = \rho_e p_e C_{p,e} T_{vap} - T_a + L_e, \text{ for } T \geq T_{vap},$$

$$q_{vap} = \rho_e p_e C_{p,e} T - T_a, \text{ for } T < T_{vap},$$

where  $n$  is the unit outward vector of  $\Gamma$ ,  $\varepsilon \in [0, 1]$  the emissivity which increases to one when the lava turns black,  $\sigma$  the Stefan-Boltzmann constant,  $T_a$  the ambient temperature (using also as the rain temperature),  $h_c$  the coefficient of heat transfer by convection which is between 5 and 25 W.m<sup>-2</sup>.K<sup>-1</sup>,  $\rho_e$  the density of water,  $C_{p,e}$  the heat capacity of water,  $L_e$  the latent heat of change of liquid-water vapor at the temperature  $T_{vap}$  and  $p_e$  the rainfall inside the Choungou Chagnoumeni crater (expressed in meters per year).

The value of 10 m.yr<sup>-1</sup> of rainfall inside the Choungou Chagnoumeni crater has been calculated thanks to the bibliographical data. The highest rainfall values recorded near the coast, at the foot of Karthala volcano is of 5888 mm/yr (western part of Nioumbadjou town; Frenken, 2005). Using an altitudinal gradient of 120 mm/100 m (Brouwers, 1973), we obtain on the caldera floor of Karthala volcano, at 2260 m a.s.l., an increase of about 2712 mm respect to the coast, reaching therefore a value of 8600 mm/yr. Considering that (1) the altitudinal gradient of 120 mm/100 m proposed by Brouwers (1973) seems to be underestimated (Mirhani, 2014) and that (2) during the highest rainfall events, the Choungou Chagnoumeni crater receive water (and sediments) from outside the Choungou Chagnoumeni crater boundary, due to the funnel shape of part of the external crater environment (see Fig. 2c), we therefore consider for the model the value of 10 m/yr of rainfall inside the Choungou Chagnoumeni crater; the bottom of the crater being flat. In order to avoid any kind of major uncertainties induced by the lack of rainfall data on the top of Karthala volcano, the heat flux lost by the rainfall events will be then discussed compared to the other cooling phenomenon.

The value of 15 °C of ambient temperature  $T_a$  has been calculated as an annual average of the Karthala temperature provided by the National Centers for Environmental Information (<https://www.ncdc.noaa.gov/>).

A condition of symmetry is given along the axis of symmetry and the ambient temperature is imposed along the external boundaries of the substratum (Fig. 4). The initial temperature on  $\Omega$  is given by a temperature  $T_0$  in the lava domain and an ambient temperature  $T_a$  in the substratum. The Table 1 resumes all the physical values used for computation.

### 3. Results

#### 3.1. ERT and CO<sub>2</sub> results

We first begin to comment, on a structural point of view, the CO<sub>2</sub> concentration and ERT straight profile performed from both sides of the western Choungou Chagnoumeni crater boundary, on 14th June 2009 ( $t_0 + 884$  days).

The CO<sub>2</sub> concentration profile allow to perfectly identify the limit of Choungou Chagnoumeni crater, with a sharp lateral increase in the CO<sub>2</sub> concentration, from values close to air CO<sub>2</sub> concentration (~400 ppm) up to 200,000 ppm, in <15 m (Fig. 3a). Such high CO<sub>2</sub> values, reaching >70%, display a clear degassing activity still in progress inside the Choungou Chagnoumeni crater, >2 years after the effusive activity, and a

**Table 1**

Physical values and units used for the numerical modeling.

Physical value	Value	Unit
$\rho$	2200	kg.m <sup>-3</sup>
$C_p$	1225	J.kg <sup>-1</sup> .K <sup>-1</sup>
$\lambda$	2	J.s <sup>-1</sup> .m <sup>-1</sup> .K <sup>-1</sup>
$\varepsilon$	0.95	–
$\sigma$	5.6704×10 <sup>-8</sup>	W.m <sup>-2</sup> .K <sup>-4</sup>
$h_c$	10	W.m <sup>-2</sup> .K <sup>-1</sup>
$\rho_e$	1000	kg.m <sup>-3</sup>
$C_{p,e}$	4180	J.kg <sup>-1</sup> .K <sup>-1</sup>
$L_e$	2.257×10 <sup>6</sup>	J.kg <sup>-1</sup>
$T_a$	288	K
$T_0$	1473	K
$T_{vap}$	373	K
$p_e$	10	m.yr <sup>-1</sup>
$H$	7.26	m
$r$	112.5	m
$L_{liq}$	1473	K
$L_{sol}$	1250	K
$L$	3.5×10 <sup>5</sup>	J.kg <sup>-1</sup>

good permeability between the source and the surface inside the Choungou Chagnoumeni crater. However, it is not possible to say if the degassing source is related to the cooling of the lava flow or to a deeper magmatic source. It is also interesting to note that at the boundary of Choungou Chagnoumeni crater and with a radius of influence of 8 m the CO<sub>2</sub> values are decreasing, with a minimum of CO<sub>2</sub> concentration located just on the boundary of Choungou Chagnoumeni crater (Fig. 3a). This decreasing of CO<sub>2</sub> concentration is probably related with higher permeability values located on the structural boundary of the crater, as we will discuss after.

About ERT results, the structural boundary with a right angle, formed by the bottom and the edge of the Choungou Chagnoumeni pit crater (thick black line in Fig. 3b), is associated only partially with a transition of resistivity. At about 7 m depth, a lateral resistivity transition exist; with values in the range 1800–16,000 Ohm.m (green-blue color) outside the Choungou Chagnoumeni crater, and with values in the range 1300–150 Ohm.m (yellow-orange-red color) inside the Choungou Chagnoumeni crater. It is very important to note that this lateral resistivity transition reach the depth of about 20 m (limit depth of the tomography), but 1) it is a gradual transition and not a drastic one, and 2) it do not reach the surface. Indeed, between 0 and about 3 m depth, there is no significant lateral transition in resistivity. That means that the petrological changes between the inner and the external part of the Choungou Chagnoumeni crater (refilled or not by the recent lava flow of January 13th 2007) is not directly involved in this lateral change of resistivity. We will talk about this topic in the Discussion section.

#### 3.2. Numerical results

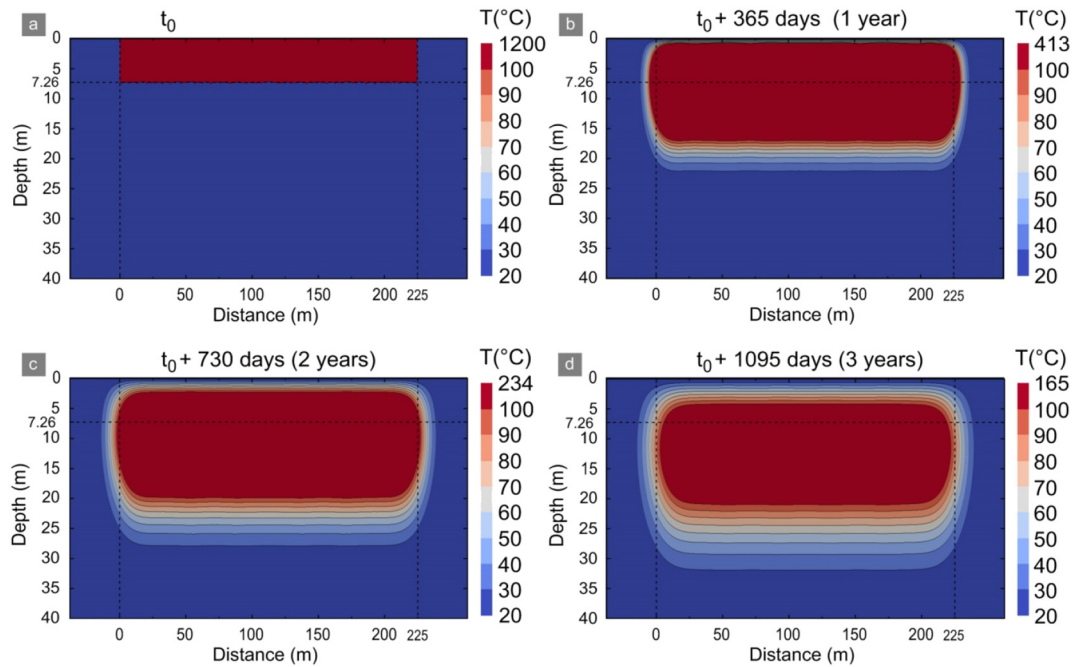
The results of the thermal modeling are shown in Fig. 5. Four different steps of cooling have been chosen:

- (1) The initial thermal conditions ( $t_0$ : Fig. 5a),
- (2) The thermal propagation after 1 year ( $t_0 + 365$  days: Fig. 5b),
- (3) The thermal propagation after 2 years ( $t_0 + 730$  days: Fig. 5c),
- (4) The thermal propagation after 3 years ( $t_0 + 1095$  days: Fig. 5d).

It is interesting to precise that from an initial condition of liquid lava with a homogeneous temperature of 1200 °C ( $T_0$ ), the maximum of temperature remaining, with the temperature modeling is of:

- (1) 413 °C at 7.26 m depth after 1 year of magma cooling,
- (2) 234 °C at 10.16 m depth after 2 years of magma cooling, and
- (3) 165 °C at 12.14 m depth after 3 years of magma cooling.





**Fig. 5.** 2D results obtained with the numerical modeling for the temperature of the lava flow cooling inside the *Choungou Chagnoumeni* crater (a) at  $t_0$ ; (b) at  $t_0 + 1$  year; (c) at  $t_0 + 2$  years; (d) at  $t_0 + 3$  years.

For readability reasons, we decided to represent in the four figures shown in Fig. 4 only the isotherms of 30, 40, 50, 60, 70, 80, 90 and 100 °C, considered to be the most interesting for our study.

In Fig. 6b, we represent the temperature of the 11 sensors located at 30 cm depth and spaced 1 m apart (see location in Fig. 2c) collected by the Campbell Scientific datalogger during the four monitoring periods, superimposed with the modeled temperature and extracted from 11 points, located at the same place of the temperature sensors.

Data collected from our CR1000 display sharp temperature variation in time for the 11 sensors. These variations are mainly related with raining events (Gaudin et al., 2015, 2017). Although these sharp temperature variations measured are not modeled, because the raining cooling impact ( $\sim 10$  m/yr) is integrated in the model over all the year, it is interesting to note that the trend of cooling modeled is in good accordance with the temperature data monitored in the field.

For the first and the fourth monitored period, the temperature data from the sensors are significantly higher than the temperature modeled, which is not the case for the second and third monitored period which is better matching with the modeled data. This shift between the data and the modeled temperature could be attributed to the raining events. Indeed, for the Comoros archipelago and more specifically the *Grande Comore* island, the raining season encompasses December up to April (<https://climate-data.org>); the first and the fourth monitored periods are both located, in part, during the raining season, while the second and the third monitored period are clearly registering data only during the dry season. During strong raining events, the huge input in water infiltration inside the ground is responsible for the heating of the soil at shallow depth due to the vaporization of the raining water at few meter depth (Barde-Cabusson et al., 2009b). Conversely, in the model, the precipitations have few effects on the global thermal flux balance. Its influence, are always  $<1\%$ , as shown on Fig. 7, which represents the contribution of each surface thermal flux.

The differences between the data and the numerical code are related to different parameters. The shallow depth of the sensors (30 cm depth) exposes the sensors to higher fluctuations forced by external parameters (e.g., strong raining events in a short period of

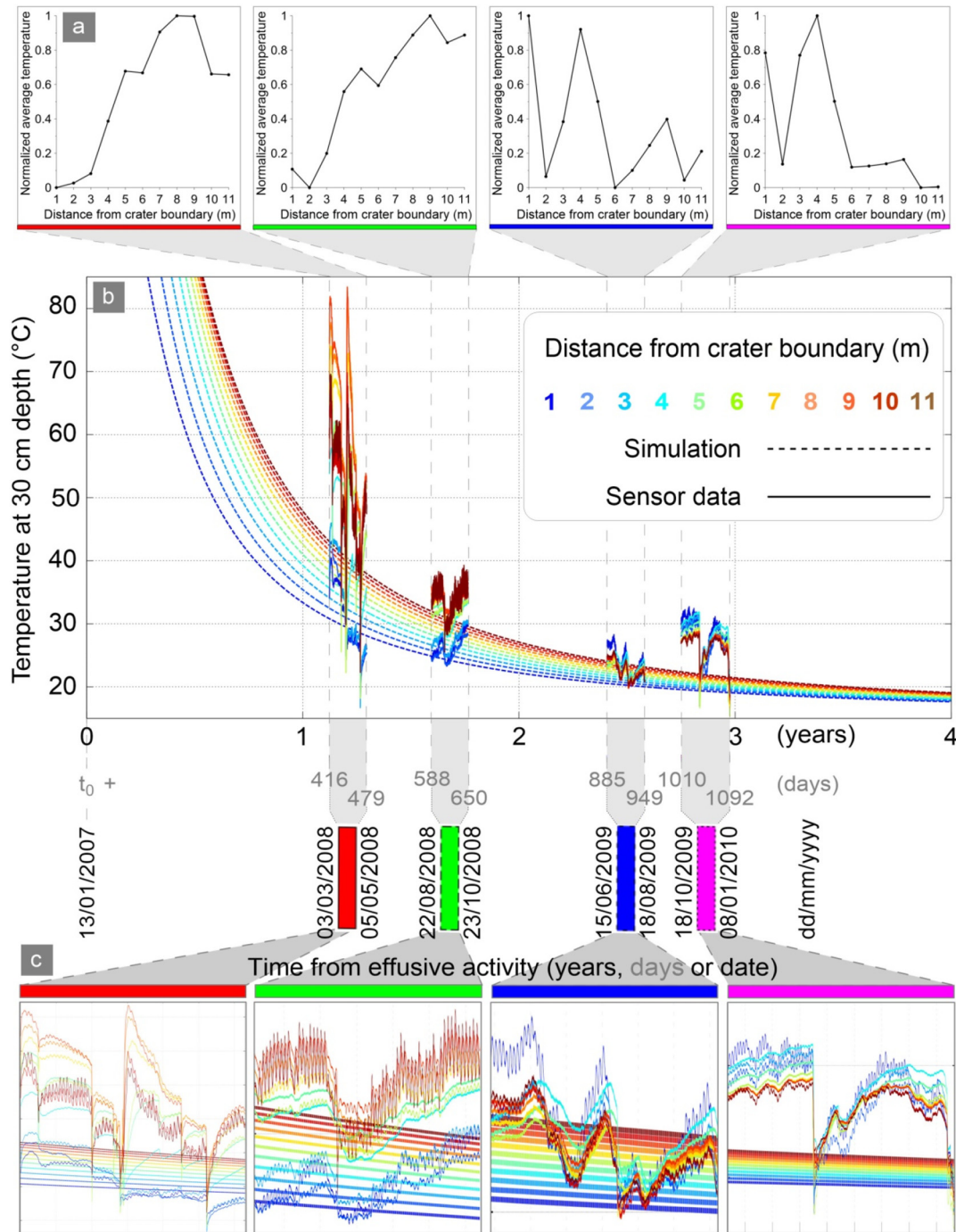
time not considered in the modeling). Moreover, the presence of sensors close to a structural boundary allows determining convective phenomenon, with rising of vapor, due to variation of lateral permeability, as it will be discussed later, but not considered in the numerical code. These aspects can explain the discrepancy between the field data and the model.

About the data monitored in the field, these latter display a surprising result: During the two first periods of monitoring ( $t_0 + 416$  days up to  $t_0 + 479$  days and  $t_0 + 588$  days up to  $t_0 + 650$  days), the temperature recorded a lateral gradient increasing with the distance from the crater boundary (sensors n°7 to 11 (yellow-orange-red-brown colors) are the sensor of highest temperature; see Fig. 6). This result has been already presented for the first temperature measured with these sensors (Fig. 2d). During the last two periods of monitoring ( $t_0 + 885$  days up to  $t_0 + 949$  days and  $t_0 + 1010$  days up to  $t_0 + 1092$  days), the spatial temperature gradient is reversed, and the highest temperature appear close to the crater boundary. Note that only the four first sensors, closer to the crater boundary (blue color) are affected by this inversion (see zoom in Fig. 6c). Although temporal variations (mainly forced by the strong raining events) are of higher amplitude of spatial variations, these latter are clearly significant (Fig. 6c).

This inversion of the temperature gradient have been also seen for the four periods of monitoring, taking into account for each sensor an average of all the measured temperature, normalized between 0 and 1. The inversion in the temperature gradient is also shown in Fig. 6a. Such inversion in the lateral temperature gradient is not visible on the modeled temperature and must be explained by an influence of the structural boundary not considered in the model (Fig. 6b).

Both ERT and temperature variations display surprising results; (1) in term of apparent incoherency between the spatial distribution of the resistivity values and the evident structural boundaries of the *Choungou Chagnoumeni* crater and (2) regarding the inversion in time, at more than two years from the effusive activity, in the shallow lateral temperature gradient.

These two aspects, dealing respectively with spatial and temporal problematic are discussed in the following session.



**Fig. 6.** (a) Normalized average temperature for each one of the eleven sensors, for the 4 periods of monitoring; (b) Temperature measured and modeled, at 30 cm depth, for the eleven sensors located inside the *Choungou Chagnoumeni* crater.

## 4. Discussion

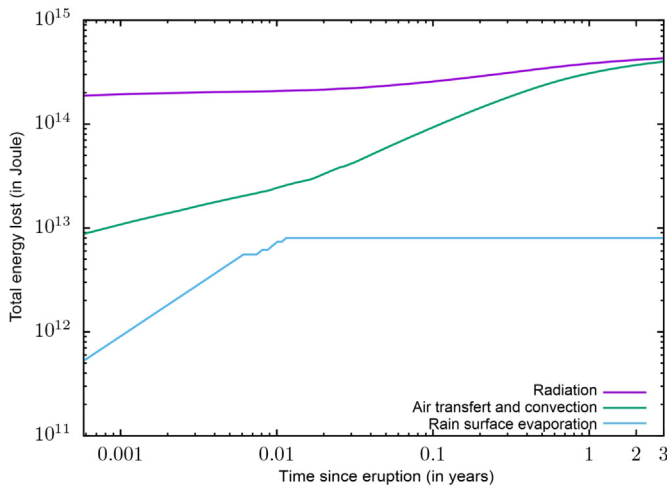
### 4.1. Spatial distribution of the resistivity values

We first focus our discussion on the ERT profile performed on 14th June 2009, so at  $t_0 + 884$  days from the effusive activity.

The most intriguing result for ERT profile is to not be able to recognize the shape of the *Choungou Chagnoumeni* crater. Indeed, several examples of crater boundaries on active volcanoes have been perfectly evidenced laterally and at depth using 2D ERT profiles (e.g. the NeoStromboli crater boundary on Stromboli volcano in [Finizola et al., 2006](#) and [Revil et al., 2011](#); or the Gran cratere on La Fossa cone of Vulcano island in [Revil et al., 2008](#) and [Barbe-Cabusson et al., 2009](#)).

In order (1) to understand the meaning of the ERT profile obtained through the western border of the *Choungou Chagnoumeni* crater and (2) to know the values of resistivity corresponding to a transition in the ERT profile, we treated the true resistivity data with the probability plot technique ([Sinclair, 1974](#)). This statistical data processing, frequently used for diffuse degassing mapping to know how many sources are responsible at depth of the degassing values measured at the surface ([Chiodini et al., 1998, 2001](#)), has been also used on ERT profiles to identify the geological boundary between pre-Pliocene carbonatic bedrock and refilling plio-quaternary alluvial deposits (see supplementary material in [Pucci et al., 2016](#)). With our ERT data set, the probability plot technique ([Fig. 3c](#)) displays a typical three range of resistivity values (three asymptotic tendencies) subdivided by two points of inflection



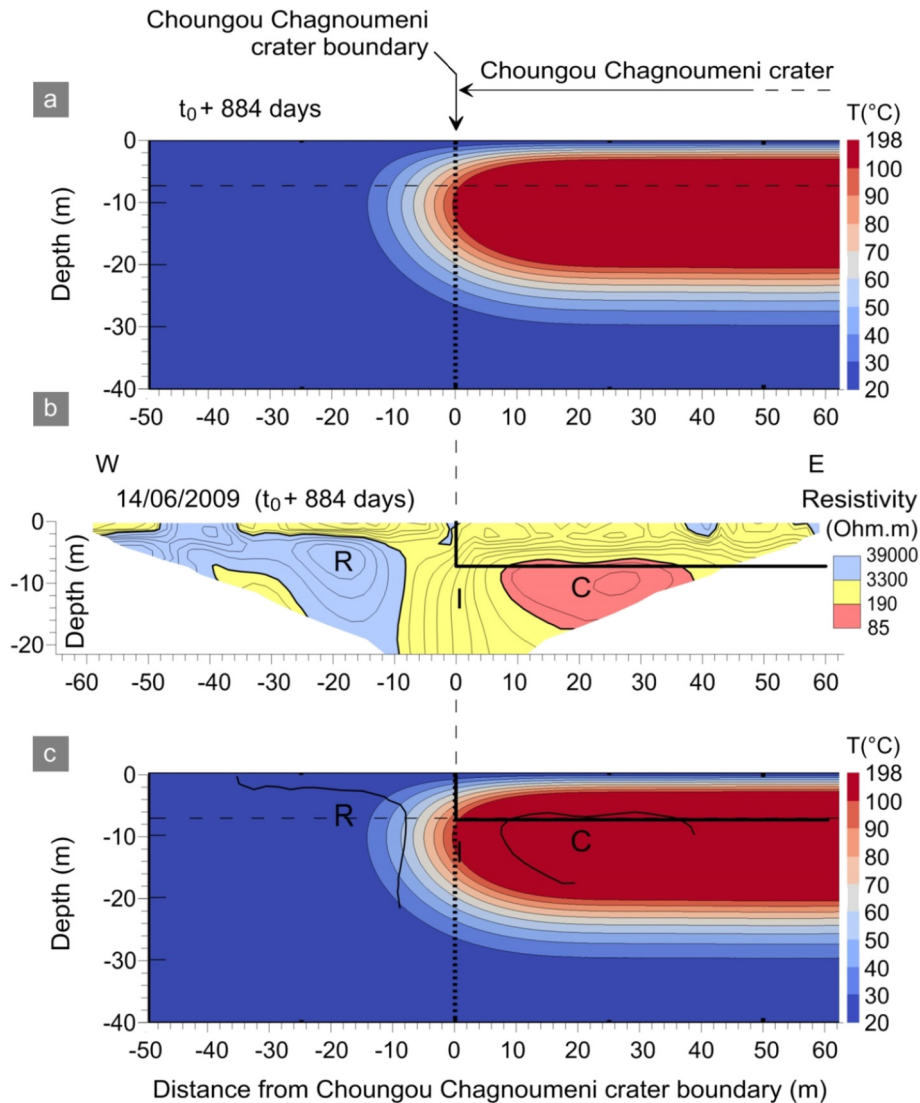


**Fig. 7.** Contribution of each energy lost (from radiation, convection at the interface soil-atmosphere and evaporation of rain) over time.

(see similar example of three lognormal populations subdivided by two inflection points; Fig. 5 in Sinclair, 1974). The two inflection points, inferred graphically, correspond to resistivity values of 3300 and 190 Ohm.m. These values correspond statistically to a transition between three groups of resistivity range: (1) lower than 190 Ohm.m., (2) between 190 and 3300 Ohm.m., and (3) higher than 3300 Ohm.m. The spatial distribution of these three ranges of resistivity values, defined from the probability plot technique is shown in Fig. 3d. Three distinct areas can be evidenced. They have been noted with the letters “C”, “I” and “R”, corresponding respectively to “Conductor”, “Intermediate” and “Resistive” body. Although the probability plot technique do not allowed understanding why the structural boundary of the *Choungou Chagnoumeni* crater is totally invisible on the ERT profile, it allowed to simply the ERT profile mainly into three areas “C”, “I” and “R”:

“C” is located below of the *Choungou Chagnoumeni* crater. The top of “C” begin at about 7 m depth and the western extension of “C” is also located at about 7 m from the western boundary of the *Choungou Chagnoumeni* crater.

From both side of the western boundary of the *Choungou Chagnoumeni* crater, from the surface up to >20 m depth, the area is



**Fig. 8.** (a) 2D results obtained with the numerical modeling for the temperature of the lava flow cooling inside the *Choungou Chagnoumeni* crater at  $t_0 + 884$  days; (b) Electrical resistivity tomography (ERT) profile showing the three ranges of resistivity values identified through the probability plot technique; (c) Superimposition of the results obtained with the numerical modeling with the two boundaries subdividing the three ranges of resistivity values. “R”, “I” and “C” stand for Resistant, Intermediate and Conductive body, respectively. Note that “C” seems to be limited to 30 m in width, only due to the lateral limit in the ERT data acquisition.

occupied by “I”. We can clearly note that the isopotential lines of “I” surround “C”, suggesting a link between the source genesis of “C” and “I”.

“R” is located only outside the western boundary of the *Choungou Chagnoumeni* crater at about 10 m from this crater. Note also the narrow vertical resistive body between the surface and about 3 m depth, located just at the boundary of the *Choungou Chagnoumeni* crater.

In order to better understand the significance of the ERT profile, we compare the ERT results with the cooling model of the 13th January 2007 lava flow that refilled the *Choungou Chagnoumeni* crater.

Using the FEM code, we stopped the temperature propagation at  $t_0 + 884$  days in order to compare the modeled temperature in the vertical section of the ground (Fig. 8a) with the distribution of the main ranges of electrical resistivity defined thanks to the probability plot technique (Fig. 8b). After  $t_0 + 884$  days, the maximum of temperature is 198 °C and it is located at 10.85 m depth. The superimposition of the boundary subdividing the conductive, intermediate and resistive body (respectively “C”, “I” and “R” in Fig. 8b) with the modeled temperature display a clear correlation (Fig. 8c); the boundary of “C” is centered inside the isotherm 100 °C, and the boundary of “R” seems to coincide with the lateral influence of the thermal propagation, between the isotherms 40 and 50 °C.

Therefore, we have to explain the good correlation between the highest temperature area (>100–200 °C) and the most conductive body (85–190 Ohm.m).

With experimental petrology, measurements of electrical conductivity versus temperature are usually performed for high temperature (>1000 °C), close to liquidus conditions, in order to study the impact of partial crystallization on electrical conductivity (Scarlato et al., 2004; with basaltic rocks of Etna; Gaillard and Marziano, 2005; with samples of basaltic composition; Poe et al., 2008; with phonotephrite of Mt. Vesuvius). In both cases, the increase of temperature is always related with an increase of several orders of magnitude in conductivity.

The presence of water inside hydrous melts constitutes also another way to increase the conductivity, and the influence of water is more marked at low (350–675 °C) than high (675–1325 °C) temperature (Pommier et al., 2008).

For temperature values closer from the temperature modeled for the *Choungou Chagnoumeni* lava flow, 884 days from the eruptive event (up to 198 °C), dry-silicate composition displays between 200 and 100 °C a resistivity range from  $10^8$  up to  $10^{14}$  Ohm.m respectively (Pommier and Le-Trong, 2011). These values are very far away from the relatively low resistivity values measured in the field (85–190 Ohm.m). Therefore, the contribution of water to reduce the resistivity values needs to be taken into account.

For water saturated rocks, the resistivity values for dacitic and andesitic tuffs are between 30 and 428 Ohm.m with aqueous solution of KCl, 70 Ohm.m at 20 °C (Llera et al., 1990). Although, in our case study, the rocks are never saturated, and the resistivity has to be higher, another parameter allows reducing the resistivity is the water temperature. Indeed, in the range of 10–200 °C, the resistivity decrease when temperature increase, and this decrease of resistivity reach two orders of magnitude between 10 and 200 °C (Light, 1984; Light et al., 2005). The consideration of volcanic rock interconnected porosity in basalts (Al-Harthi et al., 1999; Sruoga et al., 2004), typical water conductivity and water content, and water temperature reaching up to 198 °C allow to calculate with the literature cited above a range of resistivity in good accordance with our field data (between 85 and 190 Ohm.m). About the presence of water inside the *Choungou Chagnoumeni* crater, this latter can come from 1) mainly the raining events (~6–10 m/yr at the summit) and also 2) from the preferential degassing way of the magma. This aspect has been already attested by the very high soil CO<sub>2</sub> concentration associated with the profile crossing the *Choungou Chagnoumeni* crater (Fig. 3a).

In conclusion, the slow cooling of the January 13th 2007 lava flow induce the presence of a residual heat, migrating in time, which maintain relatively high temperature (100–200 °C) for rocks and fluids. This

parameter have a first order impact on the resistivity values measured, inducing low resistivity values associated with hot rock and fluid areas.

#### 4.2. Inversion in the shallow lateral temperature gradient

The second part of the discussion has been orientated on the surprising result shown in Fig. 6b and c, with the inversion in time of lateral gradient measured at 30 cm depth. This inversion occurred between the second and the third monitoring period, so between  $t_0 + 650$  days and  $t_0 + 885$  days.

This inversion is not in accordance with the temperature modeled with our code, and can be explained only considering a convective heat transfer superimposed on the conductive heat transfer.

In order to validate the notion of convection and of heat moving faster than conductive phenomenon (thanks to mass transfer), we also analyzed, for each of the eleven heat temperature sensors, the average temperature variation per day. The basis of the reflection is the following. Considering that all the temperature sensors are located precisely at the same depth and covered by the same volcanic products, all the sensors will be influenced in the same way by all kind of external parameters (e.g. diurnal variation, raining events, or wind). The sensors affected by higher amplitude variations per day would be influenced by internal parameter with a relatively quick spatial displacement and with an origin located below the sensors. Only a lateral variation of permeability associated with preferential hydrothermal fluid flow could explain higher amplitude temperature variation (energy exchange) per day.

In order to visualize such possible phenomenon, one average value was calculated for each sensor, during each one of the four periods of monitoring (Fig. 9). The results show clearly for the two first monitoring periods (red and green in Fig. 9) that the higher variations in the temperature per day are located toward the middle of the crater (sensors 5 up to 11), whereas for the two following periods of monitoring (blue and magenta in Fig. 9) the crater boundary of the *Choungou*

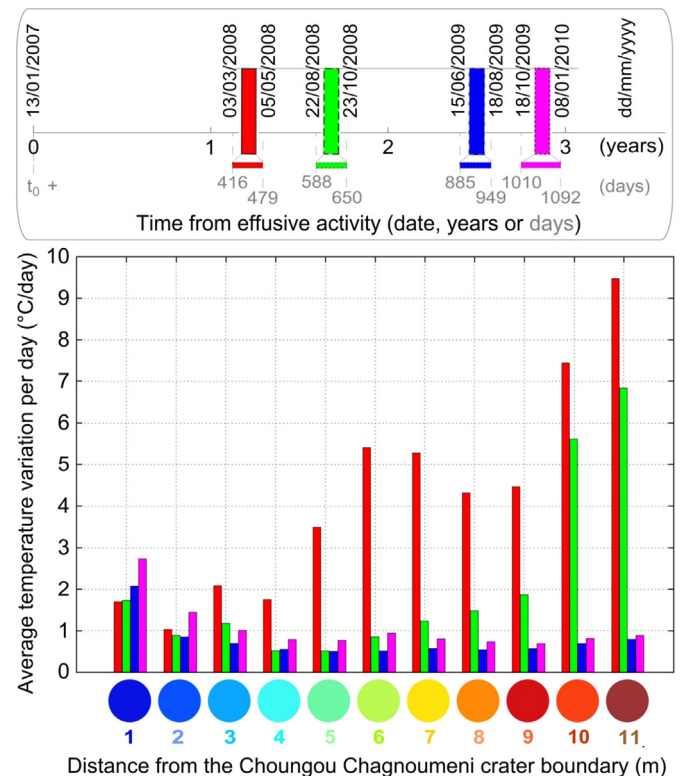


Fig. 9. Average temperature variation per day, for each one of the eleven temperature sensors, and for each one of the four monitoring periods.

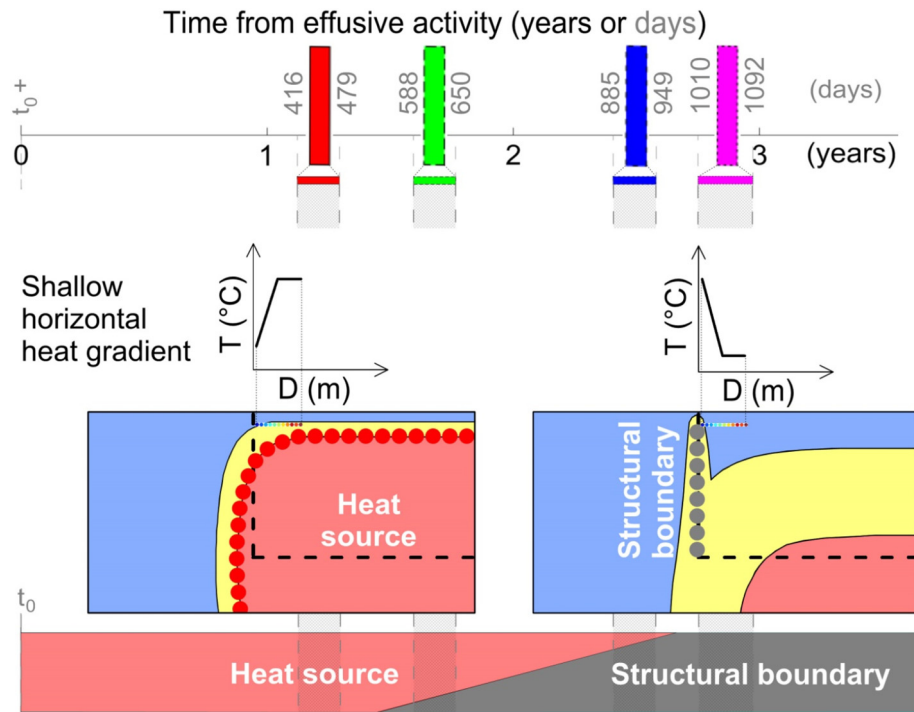


Fig. 10. Conceptual model with time evolution of the main parameter impacting the heat transfer at 30 cm depth on a crater boundary refilled by a lava flow.

Chagnoumeni crater (sensor 1) became the area of higher energetic exchange per day, for which a higher permeability have to be considered close to the crater boundary.

This interpretation of change in lateral permeability along the structural boundary of Choungou Chagnoumeni crater can be closely associated with the decrease in  $\text{CO}_2$  concentration along a distance of 8 m from both part of the crater boundary. Such a local decrease of  $\text{CO}_2$  can be also explained with an increase of permeability, allowing to the air a preferential way to enter inside the hydrothermal system. Other examples of air circulation inside a crater have been studied in the literature (Antoine et al., 2009, 2017). Note that this hypothesis of higher air content along the crater boundary is also supported by the narrow vertical resistant anomaly up to 3 m depth located exactly on the crater rim.

The lateral lithological and permeability transition bordering the crater rim have not been integrated into the model and can explain some spatial discrepancy between our field data and the model.

The concept of structural boundary (such as crater or caldera boundaries) dragging preferentially hydrothermal fluids toward the surface have been described by different authors coupling geoelectricity, temperature and diffuse degassing (Finizola et al., 2002, 2003, 2006, 2009; Revil et al., 2004, 2008, 2011; Barde-Cabusson et al., 2009a, 2009b; Carapezza et al., 2009; Bennati et al., 2011; Peltier et al., 2012).

The originality in this study have been to pointed out, at shallow depth (30 cm), the transition between an initial heat regime, mainly governed by the location of the heat source and a second heat regime, where permeability of structural boundaries begin to play a major role for the shallow heat transfer.

In our study, this transition occurred at 30 cm depth, approximately 2 years after the effusive event, but it is interesting to note that, at depth, the influence of the heat source will remain for a longer time, as shown for instance in the first order information of the ERT profile, performed at the beginning of the third monitoring period ( $t_0 + 884$  days).

The inversion in time in the shallow lateral temperature gradient, (1) from a period dominated by the importance of the heat source in the convective heat transfer to (2) a period dominated by the importance of a structural boundary of higher permeability has been schematize in Fig. 10.

The coupling of the ERT, temperature monitoring and temperature modeling techniques allow us to better understand the spatio-temporal evolution related to lava flow cooling.

## 5. Conclusion

The unusual simplicity in the cylindrical geometry of the lava flow which refilled in January 13th 2007 the Choungou Chagnoumeni crater in the summit part of Karthala volcano, allow a special study of spatio-temporal evolution of temperature and fluid flow through a new thermo-lithological boundary.

The multidisciplinary approach associating ERT measurements displayed, 884 days after this effusive event, the spatial convergence between the lowest resistivity measured and the hottest area modeled. According to the good correlation obtained in this work, it appears clearly that for hot temperature boundaries (100–200 °C), temperature constitutes the main parameter influencing resistivity data.

The lateral permeability transition related to the crater boundary was attested by the lower  $\text{CO}_2$  value associated with air contamination along the soil diffuse degassing profile. This survey also revealed the importance in structural boundaries, such as crater rims, in constituting preferential way for dragging convective mass and heat transfer, up to reverse in time the temperature gradient. This latter is influenced firstly by the hot lava, and after a period of cooling, structural boundary and the highest permeability area become the first parameter influencing the temperature gradient and fluctuation in time.

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installing the ERT profile, 2 m spacing, crossing the Chagnoumeni crater, in June 2009. CO<sub>2</sub> concentration measurements were performed thanks to the IR portable spectrometer built by Fabio Di Gangi (INGV-Palermo). The March 2008 field campaign and scientific monitoring equipment installed on the top of Karthala was funded by the PNUD (*Programme des Nations Unies pour le Développement*) program on Karthala volcano, and the June 2009 field campaign was funded by the *Conseil Régional de La Réunion* (“CORECOM” project FRROI: *Fédération des Recherches Réunion, Océan Indien*) through the University of La Réunion and *Laboratoire Géosciences Réunion-IPGP*. NB was funded by the Agence Nationale de la Recherche (ANR: <http://www.agence-nationale-recherche.fr>) through project ANR-LAVA (ANR Program: DS0902 2016; Project: ANR-16 CE39-0009, Link: [www.agence-nationale-recherche.fr/?Project=ANR-16-CE39-0009](http://www.agence-nationale-recherche.fr/?Project=ANR-16-CE39-0009)). This is ANR-LAVA contribution no. 6.

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